

PE-46 The Design of the Lynx X-ray Microcalorimeter

Abstract:

Lynx is an x-ray telescope, one of four large satellite mission concepts currently being studied by NASA to be the next astrophysics flagship mission after WFIRST. One of Lynx's three instruments is an imaging spectrometer consisting of an x-ray microcalorimeter behind an X-ray optic with an angular resolution of 0.5 arc-seconds and approximately 3 m² of area at 1 keV. This instrument will provide unparalleled diagnostics of distant extended structures and in particular will allow the detailed study of the role of cosmic feedback in the evolution of the Universe. We discuss the design and read-out of the of the array configuration including a number of sub-array options for increasing the capabilities to maximize the scientific return of the Lynx observatory.

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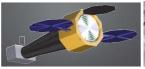
NIST/Boulder:

Douglas Bennett, Vincent Kotsubo, Benjamin Mates, Daniel Swetz, Joel Ullom

Kent Irwin, Stephen Kuenster - Stanford University Dan McCammon - University of Wisconsin Enectali Figueroa-Feliciano - Northwestern University

Benjamin Zeiger – Luxel Jeffrey Olson - Lockheed Martin Douglas Swartz – NASA/MSFC

Lynx: Revealing the Invisible Universe





Lynx: Many Science Drivers:

- · First Black Holes in the Universe and their co-evolution with
- · Cycles of (Hot) Baryons in and out of Galaxies
- · Feedback from Stars, Supernovae, and Black Holes, in all settings
- · X-ray counterparts of GW events and multi-wavelength phenomena
- · Physics of Accretion and Compact Objects



A symbol of great insight

nature of things.

Ability to see through rocks

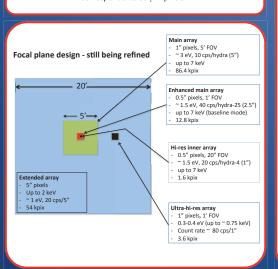
and trees to reveal the true



- The historic Accademia dei Lincei (Academy of the Lynx)
- Name based on ability to perform incisive and penetrating investigations of the natural world.
- Galileo was a proud member
- Coined the term telescope for marvelous device for peering into the cosmos.

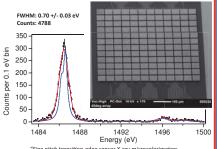
Suggested initial Lynx microcalorimeter requirements from earlier study:

- Field-of-View: At least 5' x 5'
 - 5' field-of-view with 1" pixels =>300 x 300 array => 90,000 pixels
- Energy resolution [FWHM]: < 5 eV
- Count rate capability: < 1 count per second per pixel
- For a focal length of optic of 10 m,
 - 1" corresponds to 50 μm pixels



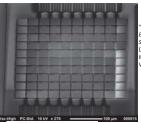
Telescope layout: **Cryostat layout:** Notional focal plane assembly layout:

Small-pixel designs 12 x 12 array on 35 µm pitch:

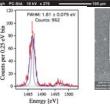


with sub-eV energy resolution at 1.5 keV", S. J. Lee et al., App. Phys. Lett. **107**, 223503 (2015)

10 x 10 array on 35 μm pitch:



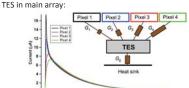
atesman et al. /ol 27, (4), (2017).



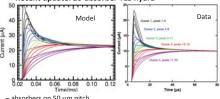


7 μm TES

Extensive use of hydra concept with up to 25 absorbers per



Recent update: 20-absorber TES hydra



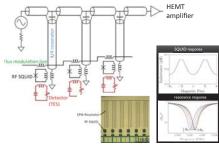
- absorbers on 50 μm pitch
- 3.4 eV demonstrated at 5.4 keV
- see poster P-59: S. Smith et al.

8x8 array of 4x5 hydras = 1280 pixels





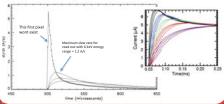
Microwave (GHz) SQUID Resonators are advancing rapidly, have the most potential, and the baseline read-out for Lynx calorimeter



- NIST has demonstrated a slew rate of 3A/s with spacing with a resonator design sufficient to read out pixels with a frequency spacing of ~ 10 MHz. See presentation O-53: Mates et al
- GSEC has demonstrated microwave SOUID read-out of X-ray microcalorimeters with almost no energy resolution degradation. - See poster PB-28 Yoon et al.

Optimization of hydra design for read-out

- Introduction of optimal Nyquist inductor to minimize slew rate & maintain position discrimination.
- Assume first pixel is removed, and its heat capacity added to TES sensor



Conclusions/outlook

- A new instrument has been designed representing a new generation in scale of X-ray microcalorimeter arrays
- Over 100,000 pixels
- New wiring schemes suitable for these arrays are being developed.
- Design exists for flight implementation of microwave SQUID read-out, assuming microwave read-out slew-rate capabilities already demonstrated.